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*L. Barletti, G. Borgioli, G. Frosali**Dipartimento di Matematica e Informatica “U.Dini”, Firenze, Italy*luigi.barletti@unifi.it, giovanni.borgioli@unifi.it,  
giovanni.frosali@unifi.it

## Semiclassical hydrodynamics of a quantum Kane model for semiconductors

In this paper we derive a semiclassical hydrodynamic system for electron densities and currents in the two energy bands of a semiconductor. We use the semiclassical Wigner equation with a  $k \cdot p$  Hamiltonian and a BGK dissipative term to construct the first two moment equations. The closure of the moment system is obtained using the Maximum Entropy Principle, by minimizing a Gibbs free-energy functional under suitable constraints. We prove that the constraint equations can be uniquely solved, i.e. that the local equilibrium state can be parametrized by the density and velocity field. Some BGK-like models are proposed to mimic the quantum interband migration.

### 1 Introduction

Description of the charge carriers dynamics in semiconductor devices is certainly a severe task, especially if one wishes to keep together a rigorous (and complete, whenever possible) physical picture with a final result (set of equations) simple enough for the numerical implementation. Hydrodynamic approach is an excellent compromise between the two requirements. Our aim is the construction of hydrodynamic equations for the electron dynamics, by means of moment method, starting from the pseudo-kinetic formulation of quantum mechanics in terms of Wigner

functions. The physical framework adopted in this paper is based on the so called  $k \cdot p$  method, [9, 13], a simple model for the description of charge transport in a semiconductor with two available energy bands.

The  $k \cdot p$  Hamiltonian has been widely studied and employed in literature (see for instance the review [5]). In particular, it has been exploited in [3, 4] to derive a semi-classical two-band diffusive model, with weak or strong external fields.

The rigorous derivation of the  $k \cdot p$  Hamiltonian from the complete Hamiltonian of an electron in a periodic potential, under a suitable homogenization scaling, is based on the concept of envelope functions and can be found in [2]. The result is a  $2 \times 2$  matrix Hamiltonian, which means that electrons in the  $k \cdot p$  description are pseudo-spinors (the pseudo-spin being related to the two energy bands). A fully-quantum treatment based on the  $k \cdot p$  method leads to non-parabolic intraband dynamics as well as to interband quantum transitions.

However, in the present semiclassical treatment, the latter aspect is lost. Nevertheless, the non-parabolic dynamics is still present and leads to non-trivial fluid models.

The semiclassical kinetic equations, that we need to get the hydrodynamic model, can be naturally expressed in terms of Wigner functions, describing statistical states of electrons in terms of quasi-distributions in phase-space. Due to pseudo-spin, the standard scalar Wigner function has to be substituted by a matrix-valued Wigner function. Such a matrix can be projected on the two energy subspaces, thus obtaining two distributions of electrons, corresponding to the two energy bands. Then, the macroscopic fluid quantities can be obtained by taking moments of the band-projected Wigner function, which have the physical meaning of densities  $n_{\pm}$  and velocity field  $\mathbf{u}_{\pm}$ , where the subscript  $\pm$  means  $+$ , the upper band, and  $-$ , the lower band (see Eqs. (24) and (25)). The Wigner formalism, moreover, permits the introduction of a well justified BGK term (see [1, 8]) which takes in account the interaction phenomena leading to a local equilibrium relaxation. Thanks to this relaxation mechanism we can assume that, in a time-scale larger than the relaxation time, the system is in a local equilibrium state. The latter is chosen according to the Maximum Entropy Principle (MEP), i.e. as the most probable microscopic state, given the observed macroscopic moments  $n_{\pm}$  and  $\mathbf{u}_{\pm}$ . This strategy, as usual, provides a closure of the moment equations.

The paper is organized in the following way: in section 2 we present the

$k \cdot p$  Hamiltonian. The presence of the two bands is treated introducing a pseudo-spinorial formulation via a representation on the Pauli matrices basis. In section 3 we deduce the Wigner-BGK equations for our model. The Wigner matrix is decomposed in its scalar part  $w_0$  and its pseudo-spinorial part  $\vec{w}$ .  $\vec{w}$  is further split in a part parallel to the direction of the pseudo-spinorial part of the Hamiltonian,  $w_S$ , and a part orthogonal to it,  $\vec{w}_\perp$ . This representation discovers itself useful in the evaluation of the moments for the Wigner equation, since the contribution of  $\vec{w}_\perp$  vanishes. In Section 4 we deduce the moment equations of zeroth and first order, where appear the tensors  $\mathbb{P}_\pm$  and  $\mathbb{Q}_\pm$ , which can be interpreted as the pressure and effective-mass tensors. In Section 5 the application of the MEP implies that these tensors depend on two Lagrange multipliers, a scalar one,  $A_\pm$ , and a vector one,  $\mathbf{B}_\pm$ . The closure of the moment equations requests the study of the dependence of the tensors on the macroscopic quantities,  $n_\pm$ , the numerical density and  $\mathbf{u}_\pm$ , the velocity field. In Theorem 1 we prove that  $\mathbf{B}_\pm$  (and  $A_\pm$ , as a consequence) is a smooth globally invertible function of the macroscopic quantities.

Since in semiclassical limit the quantum interference terms between the two bands disappear, in Section 6 we examine some models that enable the reintroduction of this aspect. We propose there three different BGK-like terms which satisfy this condition.

## 2 The $k \cdot p$ model

The simplest possible description of an electron in a semiconductor crystal with two energy bands (e. g. “valence” and “conduction”) is obtained from a periodic Hamiltonian by means of the  $k \cdot p$  method [9, 13] and consists of a  $2 \times 2$  Hamiltonian of the following form:

$$H = \begin{pmatrix} -\frac{\hbar^2}{2m}\Delta + E_g/2 & -\frac{\hbar^2}{m}\mathbf{K} \cdot \nabla \\ \frac{\hbar^2}{m}\mathbf{K} \cdot \nabla & -\frac{\hbar^2}{2m}\Delta - E_g/2 \end{pmatrix}. \quad (1)$$

Here,  $E_g$  is the band-gap and  $\mathbf{K} = (K_1, K_2, K_3)$  is the matrix element of the gradient operator between the Bloch functions  $b_\pm$  of the upper (+) and lower (−) bands, evaluated at zero pseudo-momentum:

$$\mathbf{K} = \int_{\text{lattice cell}} \bar{b}_+(\mathbf{x}) \nabla b_-(\mathbf{x}) d\mathbf{x},$$

$\hbar$  is Planck's constant over  $2\pi$  and  $m$  is the electron mass. The  $k \cdot p$  model has to be completed by adding an "external" potential term  $qV$  (where  $q > 0$  denotes the elementary charge), accounting for all electric fields except the crystal one. The electric potential  $V(\mathbf{x})$  can be either fixed or self-consistently given by a Poisson equation.

The  $k \cdot p$  Hamiltonian  $H$  is the quantization of the classical matrix-valued symbol

$$h(\mathbf{p}) = \begin{pmatrix} \frac{p^2}{2m} + E_g/2 & -i\frac{\hbar}{m}\mathbf{K} \cdot \mathbf{p} \\ i\frac{\hbar}{m}\mathbf{K} \cdot \mathbf{p} & \frac{p^2}{2m} - E_g/2 \end{pmatrix}, \quad (2)$$

where  $p = |\mathbf{p}|$ .

In this paper we make the choice to decompose any  $2 \times 2$  complex matrix in the basis of Pauli matrices

$$\sigma_0 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \quad \sigma_1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma_2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma_3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix},$$

(the coefficients of the decomposition will be real if the matrix is hermitian). The operators  $\sigma_1, \sigma_2, \sigma_3$  are called "pseudo-spin components" in this context. Putting

$$\boldsymbol{\alpha} = (\alpha_1, \alpha_2, \alpha_3) := \frac{\hbar}{m}\mathbf{K} \quad \text{and} \quad \gamma := E_g/2, \quad (3)$$

we can write

$$h(\mathbf{p}) = \frac{p^2}{2m}\sigma_0 + \boldsymbol{\alpha} \cdot \mathbf{p}\sigma_2 + \gamma\sigma_3 = h_0(\mathbf{p})\sigma_0 + \vec{h}(\mathbf{p}) \cdot \vec{\sigma}, \quad (4)$$

where

$$h_0(\mathbf{p}) = \frac{p^2}{2m}, \quad \vec{h}(\mathbf{p}) = (0, \boldsymbol{\alpha} \cdot \mathbf{p}, \gamma),$$

and, as usual,  $\vec{\sigma} = (\sigma_1, \sigma_2, \sigma_3)$  is the formal vector of Pauli matrices. Here and in the following we adopt the arrow notation for three-vectors, such as  $\vec{h}(\mathbf{p})$ , that are the pseudo-spinorial part of the Pauli coefficients. Instead, we do not use the arrow notation for "cartesian" three-vectors such as  $\mathbf{x}, \mathbf{p}, \mathbf{K}, \boldsymbol{\alpha}$ , etc. The dispersion relation for the free Hamiltonian  $H$  is easily obtained by computing the ( $\mathbf{p}$ -dependent) eigencouples of the symbol  $h(\mathbf{p})$ . This yields to the energy bands

$$E_{\pm}(\mathbf{p}) = \frac{p^2}{2m} \pm \sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2} = \frac{p^2}{2m} \pm |\vec{h}(\mathbf{p})| \quad (5)$$

and to the corresponding normalized energy eigenvectors

$$\psi_{\pm}^p = \frac{1}{\sqrt{2(1 \pm \nu_3(\mathbf{p}))}} \begin{pmatrix} \nu_3(\mathbf{p}) \pm 1 \\ \nu_1(\mathbf{p}) + i\nu_2(\mathbf{p}) \end{pmatrix}, \quad (6)$$

where we have introduced

$$\vec{\nu}(\mathbf{p}) = (\nu_1(\mathbf{p}), \nu_2(\mathbf{p}), \nu_3(\mathbf{p})) = \frac{\vec{h}(\mathbf{p})}{|\vec{h}(\mathbf{p})|} = \frac{(0, \boldsymbol{\alpha} \cdot \mathbf{p}, \gamma)}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}}. \quad (7)$$

The two eigenprojections  $P_{\pm}(\mathbf{p})$ , that we call band-projections, are therefore given by

$$P_{\pm}(\mathbf{p}) = \psi_{\pm}^p \otimes \psi_{\pm}^p = \frac{1}{2}(\sigma_0 \pm \vec{\nu}(\mathbf{p}) \cdot \vec{\sigma}) \quad (8)$$

and we can clearly write

$$h(\mathbf{p}) = E_+(\mathbf{p})P_+(\mathbf{p}) + E_-(\mathbf{p})P_-(\mathbf{p}). \quad (9)$$

Important quantities associated to the energy bands are the semiclassical velocities  $\mathbf{v}_{\pm}$

$$\mathbf{v}_{\pm} = \nabla_{\mathbf{p}} E_{\pm}(\mathbf{p}) = \frac{\mathbf{p}}{m} \pm \frac{\boldsymbol{\alpha} \cdot \mathbf{p}}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \boldsymbol{\alpha} = \frac{\mathbf{p}}{m} \pm \nu_2 \boldsymbol{\alpha} \quad (10)$$

and the effective-mass tensor  $\mathbb{M}_{\pm}(\mathbf{p})$  defined by [2]

$$\mathbb{M}_{\pm}^{-1}(\mathbf{p}) = \nabla_{\mathbf{p}} \otimes \nabla_{\mathbf{p}} E_{\pm}(\mathbf{p}) = \frac{1}{m} \mathbb{I} \pm \frac{\gamma^2 \boldsymbol{\alpha} \otimes \boldsymbol{\alpha}}{((\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2)^{3/2}}. \quad (11)$$

where  $\mathbb{I}$  is the identity matrix.

### 3 Wigner-BGK equations for the $k \cdot p$ model

Let  $\rho_{ij}(\mathbf{x}, \mathbf{y}, t)$ ,  $1 \leq i, j \leq 3$ , be the density matrix describing the quantum statistical state of electrons with Hamiltonian (1). The corresponding kinetic-like description is provided by the Wigner matrix  $w_{ij}(\mathbf{x}, \mathbf{p}, t)$  defined by [14, 16, 3]

$$w_{ij}(\mathbf{x}, \mathbf{p}, t) = \frac{1}{(2\pi\hbar)^{3/2}} \int_{\mathbb{R}^3} \rho_{ij} \left( \mathbf{x} + \frac{\boldsymbol{\xi}}{2}, \mathbf{x} - \frac{\boldsymbol{\xi}}{2}, t \right) e^{-i\mathbf{p} \cdot \boldsymbol{\xi} / \hbar} d\boldsymbol{\xi}. \quad (12)$$

The Wigner matrix  $w = (w_{ij})$  is hermitian,

$$w(\mathbf{x}, \mathbf{p}, t) = w^*(\mathbf{x}, \mathbf{p}, t),$$

and, consequently, its Pauli representation

$$w = w_0 \sigma_0 + \vec{w} \cdot \vec{\sigma}, \quad \vec{w} = (w_1, w_2, w_3) \quad (13)$$

has real components  $w_k(\mathbf{x}, \mathbf{p}, t)$ ,  $0 \leq k \leq 3$ .

Considering  $P_{\pm}$  and  $\vec{v}$ , as defined in (8) and (7), the two scalar functions

$$w_{\pm} = \text{Tr}(P_{\pm} w) = w_0 \pm \vec{v} \cdot \vec{w} \quad (14)$$

can be semi-classically interpreted as the phase-space distributions of electrons in the two energy bands  $E_{\pm}$  [3] and will play a central role in the following. Moreover, if  $w_s = \vec{v} \cdot \vec{w}$ , we have the obvious relations

$$w_{\pm} = w_0 \pm w_s, \quad w_0 = \frac{w_+ + w_-}{2}, \quad w_s = \frac{w_+ - w_-}{2}, \quad (15)$$

and  $w_s$  has therefore the meaning of ‘‘band polarization’’. It will be convenient, moreover, to introduce a notation for the perpendicular part of  $\vec{w}$  with respect to  $\vec{v}$  by putting

$$\vec{w} = w_s \vec{v} + \vec{w}_{\perp}. \quad (16)$$

Assume now that the dynamics of the density matrix  $\rho(\mathbf{x}, \mathbf{y}, t)$  is given by the von Neumann equation (Schrödinger equation for mixed states)

$$i\hbar \frac{\partial \rho}{\partial t} = (H_{\mathbf{x}} - H_{\mathbf{y}})\rho + (V(\mathbf{x}) - V(\mathbf{y}))\sigma_0 \rho,$$

where  $H_{\mathbf{x}}$  and  $H_{\mathbf{y}}$  denote the  $k \cdot p$  Hamiltonian (1) acting, respectively, on the  $\mathbf{x}$  and  $\mathbf{y}$  variables, and  $V$  is an external and/or self-consistent electric field. Then, using (12) and (13), it is not difficult to prove that, up to terms of order  $\hbar^2$ , the evolution equations for the time dependent Pauli-Wigner functions are the following

$$\begin{cases} \frac{\partial w_0}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} w_0 + \mathbf{F} \cdot \nabla_{\mathbf{p}} w_0 + \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_2 = 0, \\ \frac{\partial \vec{w}}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} \vec{w} + \mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{w} + \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_0 \vec{e}_2 - \frac{2}{\hbar} \vec{h}(\mathbf{p}) \times \vec{w} = 0. \end{cases} \quad (17)$$

Here,  $\vec{h}(\mathbf{p}) = (0, \boldsymbol{\alpha} \cdot \mathbf{p}, \gamma)$ ,  $\vec{e}_2 = (0, 1, 0)$  and  $\mathbf{F} = -\nabla V$  denotes the external force corresponding to the electric potential  $V$ .

In order to supplement system (17), which describes a conservative Hamiltonian dynamics, with a collisional mechanism, we insert a BGK (Bhatnagar-Gross-Krook) collisional relaxation-time term. This term mimics the collisions that force the system towards a local equilibrium and it is characterized by the relaxation time  $\tau_c$ , which is assumed to be the same constant for all components. The system, which will be referred to as ‘‘Wigner-BGK’’ (WBGK) equations, takes the new form

$$\begin{cases} \frac{\partial w_0}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} w_0 + \mathbf{F} \cdot \nabla_{\mathbf{p}} w_0 + \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_2 = \frac{g_0 - w_0}{\tau_c}, \\ \frac{\partial \vec{w}}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} \vec{w} + \mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{w} + \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_0 \vec{e}_2 - \frac{2}{\hbar} \vec{h}(\mathbf{p}) \times \vec{w} = \frac{\vec{g} - \vec{w}}{\tau_c}, \end{cases} \quad (18)$$

where  $g = g_0 \sigma_0 + \vec{g} \cdot \vec{\sigma}$  is a local-equilibrium Wigner matrix that will be specified later on.

We now extract from Eq. (18), equations for the band distributions  $w_+$  and  $w_-$  (see definition (14)). For this purpose we introduce the orthonormal basis  $(\vec{n}_1, \vec{n}_2, \vec{v})$ , where  $\vec{n}_1 \equiv \vec{e}_1 = (1, 0, 0)$  and  $\vec{n}_2$  is chosen such that  $\vec{n}_1 \times \vec{n}_2 = \vec{v}$ . Using the decomposition  $\vec{w} = w_s \vec{v} + \vec{w}_\perp$  (see ((16)) and taking account that  $w_2 = w_s \vec{v} \cdot \vec{e}_2 + \vec{w}_\perp \cdot \vec{e}_2$ , with

$$\vec{e}_2 = \frac{\boldsymbol{\alpha} \cdot \mathbf{p}}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \vec{v} + \frac{\gamma}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \vec{n}_2, \quad (19)$$

we rewrite the first of equations (18) as

$$\begin{aligned} \frac{\partial w_0}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} w_0 + \mathbf{F} \cdot \nabla_{\mathbf{p}} w_0 + \frac{\boldsymbol{\alpha} \cdot \mathbf{p}}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_s \\ + \frac{\gamma}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \vec{n}_2 \cdot (\boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} \vec{w}_\perp) = \frac{g_0 - w_0}{\tau_c}. \end{aligned} \quad (20)$$

Concerning the second of equations (18), using again (19), we have

$$\begin{aligned} \frac{\partial}{\partial t} (w_s \vec{v} + \vec{w}_\perp) + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} (w_s \vec{v} + \vec{w}_\perp) + \frac{\boldsymbol{\alpha} \cdot \mathbf{p} \vec{v} + \gamma \vec{n}_2}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_0 \\ + \vec{v} \mathbf{F} \cdot \nabla_{\mathbf{p}} w_s + (\mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{v}) w_s + \mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{w}_\perp \\ = \frac{2}{\hbar} \vec{h}(\mathbf{p}) \times \vec{w}_\perp + \frac{g_s - w_s}{\tau_c} \vec{v} + \frac{\vec{g}_\perp - \vec{w}_\perp}{\tau_c}. \end{aligned}$$

Decomposing this equation in the parallel and perpendicular parts with respect to  $\vec{\nu}$ , and using  $\vec{\nu} \cdot (\mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{\nu}) = 0$ , we obtain an equation for  $w_s$ :

$$\begin{aligned} \frac{\partial w_s}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} w_s + \mathbf{F} \cdot \nabla_{\mathbf{p}} w_s + \frac{\boldsymbol{\alpha} \cdot \mathbf{p}}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_0 \\ + \vec{\nu} \cdot (\mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{w}_{\perp}) = \frac{g_s - w_s}{\tau_c}, \end{aligned} \quad (21)$$

and an equation for  $\vec{w}_{\perp}$ :

$$\begin{aligned} \frac{\partial \vec{w}_{\perp}}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} \vec{w}_{\perp} + (\mathbf{F} \cdot \nabla_{\mathbf{p}}) w_s + (\mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{w}_{\perp})_{\perp} \\ + \frac{\gamma \vec{n}_2}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_0 = \frac{2}{\hbar} \vec{h}(\mathbf{p}) \times \vec{w}_{\perp} + \frac{\vec{g}_{\perp} - \vec{w}_{\perp}}{\tau_c}, \end{aligned} \quad (22)$$

(which will not be used in the following). Then, recalling (15) and (10), equations for  $w_+$  and  $w_-$  are now readily obtained from (20) and (21):

$$\begin{aligned} \frac{\partial w_{\pm}}{\partial t} + \mathbf{v}_{\pm} \cdot \nabla_{\mathbf{x}} w_{\pm} + \mathbf{F} \cdot \nabla_{\mathbf{p}} w_{\pm} + \frac{\gamma}{\sqrt{(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2}} \vec{n}_2 \cdot (\boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} \vec{w}_{\perp}) \\ \pm \vec{\nu} \cdot (\mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{w}_{\perp}) = \frac{g_{\pm} - w_{\pm}}{\tau_c}. \end{aligned} \quad (23)$$

## 4 Moment equations and entropy closure

The local equilibrium Wigner matrix  $g = g_0 \sigma_0 + \vec{g} \cdot \vec{\sigma}$  is given by the MEP and is, therefore, the maximizer of a suitable entropy functional (which depends on the particle statistics) under the constraint of given macroscopic moments [10, 15]. We make the following assumptions:

1. the system is in thermal equilibrium at constant temperature  $T > 0$  (e.g. with a phonon bath);
2. the electron statistics is well approximated by Maxwell-Boltzmann distribution (in the semiclassical approach);
3. the observed macroscopic moments are the densities

$$n_{\pm}(\mathbf{x}, t) = \int_{\mathbb{R}^3} w_{\pm}(\mathbf{x}, \mathbf{p}, t) d\mathbf{p} \quad (24)$$



and the velocity field

$$\mathbf{u}_{\pm}(\mathbf{x}, t) = \frac{1}{n_{\pm}(\mathbf{x}, t)} \int_{\mathbb{R}^3} \mathbf{v}_{\pm}(\mathbf{p}) w_{\pm}(\mathbf{x}, \mathbf{p}, t) d\mathbf{p} \quad (25)$$

of the electrons in the two energy bands.

It follows from the above assumptions that the local equilibrium  $g$  must be sought as the minimizer of the Gibbs free-energy functional

$$\mathcal{E}(w) = \int_{\mathbb{R}^6} \text{Tr} \{k_B T(w \log w - w) + hw\} d\mathbf{p} d\mathbf{x}, \quad (26)$$

among all positive-definite Wigner matrices  $w$  sharing the macroscopic moments (24) and (25). In (26),  $k_B$  is the Boltzmann constant,  $h$  is the matrix-valued symbol of the Hamiltonian (see (2)), and  $\log w$  is the matrix logarithm. It can be shown [3] that the solution  $g$  of such constrained minimization problem is given by

$$g_{\pm}(\mathbf{x}, \mathbf{p}, t) = e^{-\beta E_{\pm}(\mathbf{p}) + \mathbf{B}_{\pm} \cdot \mathbf{v}_{\pm}(\mathbf{p}) + A_{\pm}}, \quad \vec{g}_{\perp} = 0, \quad (27)$$

where  $\beta = (k_B T)^{-1}$ , and  $A_{\pm} = A_{\pm}(\mathbf{x}, t)$  and  $\mathbf{B}_{\pm} = \mathbf{B}_{\pm}(\mathbf{x}, t)$  are Lagrange multipliers to be determined from the constraint equations

$$\begin{aligned} \int_{\mathbb{R}^3} g_{\pm}(\mathbf{x}, \mathbf{p}, t) d\mathbf{p} &= n_{\pm}(\mathbf{x}, t), \\ \int_{\mathbb{R}^3} \mathbf{v}_{\pm}(\mathbf{p}) g_{\pm}(\mathbf{x}, \mathbf{p}, t) d\mathbf{p} &= n_{\pm}(\mathbf{x}, t) \mathbf{u}_{\pm}(\mathbf{x}, t). \end{aligned} \quad (28)$$

Let us now assume that the time-scale over which the system is observed is much larger than the relaxation time  $\tau_c$  (the so-called hydrodynamic asymptotics). In this limit, we have that  $w \rightarrow g$  and we can rewrite Eq. (23) with  $w_{\pm} = g_{\pm}$  and  $\vec{w}_{\perp} = \vec{g}_{\perp} = 0$ , obtaining that the local equilibrium function satisfies

$$\frac{\partial g_{\pm}}{\partial t} + \mathbf{v}_{\pm} \cdot \nabla_{\mathbf{x}} g_{\pm} + \mathbf{F} \cdot \nabla_{\mathbf{p}} g_{\pm} = 0. \quad (29)$$

**Remark 1** The quantum interference terms (i.e. the terms containing  $\vec{w}_{\perp}$  in Eq. (23)), which are responsible for quantum coupling between the two bands [11], have disappeared in our semiclassical hydrodynamic picture because  $\vec{g}_{\perp} = 0$ . When dealing with the semiclassical diffusive

limit, however, we have to consider terms of order  $\hbar$  in the semiclassical expansion of the quantum equilibrium (our  $g$  is the leading order of such expansion) and band-coupling interference terms appear [3, 6].  $\square$

Integrating Eq. (29) over  $\mathbb{R}^3$ , and using the constraints (28), we have

$$\frac{\partial n_{\pm}}{\partial t} + \nabla_{\mathbf{x}} \cdot (n_{\pm} \mathbf{u}_{\pm}) = 0 \quad (30)$$

that is the continuity equation for  $n_{\pm}$ . Multiplying Eq. (29) by  $\mathbf{v}_{\pm}$  and integrating over  $\mathbf{p}$ , we obtain the first-order moment equation

$$\frac{\partial(n_{\pm} \mathbf{u}_{\pm})}{\partial t} + \nabla_{\mathbf{x}} \cdot \mathbb{P}_{\pm} - \mathbf{F} \cdot \mathbb{Q}_{\pm} = 0, \quad (31)$$

that is the momentum balance equation, where the tensors  $\mathbb{P}_{\pm}$  and  $\mathbb{Q}_{\pm}$  are defined as follows:

$$\mathbb{P}_{\pm} = \int_{\mathbb{R}^3} \mathbf{v}_{\pm} \otimes \mathbf{v}_{\pm} g_{\pm} d\mathbf{p}, \quad \mathbb{Q}_{\pm} = \int_{\mathbb{R}^3} (\nabla_{\mathbf{p}} \otimes \mathbf{v}_{\pm}) g_{\pm} d\mathbf{p}. \quad (32)$$

Recalling (10) and (11), the tensor  $\mathbb{Q}_{\pm}$ , which ‘‘mediates’’ the action of the force  $\mathbf{F}$ , can be written as

$$\mathbb{Q}_{\pm} = \int_{\mathbb{R}^3} (\nabla_{\mathbf{p}} \otimes \nabla_{\mathbf{p}} E_{\pm}) g_{\pm} d\mathbf{p} = \int_{\mathbb{R}^3} \mathbb{M}_{\pm}^{-1}(\mathbf{p}) g_{\pm} d\mathbf{p}, \quad (33)$$

showing that  $\mathbb{Q}_{\pm}$  is the average inverse effective-mass. For suitable values of  $\alpha$  and  $\gamma$ ,  $\mathbb{Q}_{-}$  can be negative: in this case the lower-band electrons behave like positive-charged carriers (holes).

We remark that the functions  $g_{\pm}$  have been determined by the maximum entropy principle and depend implicitly on the moments  $n_{\pm}$  and  $\mathbf{u}_{\pm}$  because the constraints (28). In this sense, the tensors  $\mathbb{P}_{\pm}$  and  $\mathbb{Q}_{\pm}$  can be regarded as functions of  $n_{\pm}$  and  $\mathbf{u}_{\pm}$ , making the hydrodynamic system (30) + (31) formally closed.

For future reference let us summarize here the hydrodynamic model that we have obtained: it consists of the moment equations

$$\begin{cases} \frac{\partial n_{\pm}}{\partial t} + \nabla_{\mathbf{x}} \cdot (n_{\pm} \mathbf{u}_{\pm}) = 0, \\ \frac{\partial(n_{\pm} \mathbf{u}_{\pm})}{\partial t} + \nabla_{\mathbf{x}} \cdot \mathbb{P}_{\pm} - \mathbf{F} \cdot \mathbb{Q}_{\pm} = 0, \end{cases} \quad (34)$$

and of the closure relations (32) and (28).

## 5 The constraint equations

In this section we study the problem of how writing in a more explicit way the moment equations, that is expressing the Lagrange multipliers  $A$  and  $\mathbf{B}_\pm$ , and consequently the tensors  $\mathbb{P}_\pm$  and  $\mathbb{Q}_\pm$ , as functions of the moments  $n_\pm$  and  $\mathbf{u}_\pm$ .

In order to simplify the notations we note that, both in the moment equations (34) and in the constraint equations (28), the  $+$  and  $-$  quantities are completely decoupled (unless coupling mechanisms are introduced, as we will discuss in Section 6). Then, we can safely drop the  $\pm$  labels everywhere, bearing in mind, however, that the  $+$  and  $-$  problems are formally identical but physically different, because energies, velocities and effective-masses are different in the two bands.

In order to stress the dependence of the local-equilibrium on the Lagrange multipliers we put

$$\phi(A, \mathbf{B}, \mathbf{p}) = e^{-\beta E(\mathbf{p}) + \mathbf{B} \cdot \mathbf{v}(\mathbf{p}) + A}, \quad (35)$$

and rewrite the constraint equations (28) as follows:

$$\int_{\mathbb{R}^3} \phi(A, \mathbf{B}, \mathbf{p}) d\mathbf{p} = n, \quad \int_{\mathbb{R}^3} \mathbf{v}(\mathbf{p}) \phi(A, \mathbf{B}, \mathbf{p}) d\mathbf{p} = n\mathbf{u}, \quad (36)$$

(recall that we are suppressing the labels  $\pm$ , and that  $A$ ,  $\mathbf{B}$ ,  $n$  and  $\mathbf{u}$  are functions of  $(\mathbf{x}, t)$ ). Equations (36) have to be regarded as a system of four scalar equations in the unknowns  $A$  and  $\mathbf{B} = (B_1, B_2, B_3)$ , for given  $n > 0$  and  $\mathbf{u} = (u_1, u_2, u_3) \in \mathbb{R}^3$ .

Let us introduce the function  $f(\mathbf{B})$  defined by

$$e^{f(\mathbf{B})} = \int_{\mathbb{R}^3} e^{-\beta E(\mathbf{p}) + \mathbf{B} \cdot \mathbf{v}(\mathbf{p})} d\mathbf{p}. \quad (37)$$

By using

$$\mathbf{v}(\mathbf{p}) \phi(A, \mathbf{B}, \mathbf{p}) = \nabla_{\mathbf{B}} \phi(A, \mathbf{B}, \mathbf{p}),$$

we obtain that the constraint system (36) is (formally) equivalent to

$$\begin{cases} e^A e^{f(\mathbf{B})} = n, \\ \nabla_{\mathbf{B}} f(\mathbf{B}) = \mathbf{u}. \end{cases} \quad (38)$$

From Eq. (38) we see that  $\mathbf{B}$  only depends on  $\mathbf{u}$  and, once  $\mathbf{B}$  is solved from the second equation as function of  $\mathbf{u}$ , the remaining unknown  $A$  is determined by  $e^A = n e^{-f(\mathbf{B})}$ . Moreover, using

$$\mathbf{v}(\mathbf{p}) \otimes \mathbf{v}(\mathbf{p}) \phi(A, \mathbf{B}, \mathbf{p}) = \nabla_{\mathbf{B}} \otimes (\nabla_{\mathbf{B}} \phi(A, \mathbf{B}, \mathbf{p})),$$

the tensor  $\mathbb{P}$  (see definition (32)) can be written as

$$\begin{aligned} \mathbb{P} &= e^A \int_{\mathbb{R}^3} \nabla_{\mathbf{B}} \otimes \left( \nabla_{\mathbf{B}} e^{-\beta E(\mathbf{p}) + \mathbf{B} \cdot \mathbf{v}(\mathbf{p})} \right) d\mathbf{p} \\ &= e^A \nabla_{\mathbf{B}} \otimes \left( \nabla_{\mathbf{B}} e^{f(\mathbf{B})} \right) = e^A \nabla_{\mathbf{B}} \otimes \left( e^{f(\mathbf{B})} (\nabla_{\mathbf{B}} f(\mathbf{B})) \right) \\ &= e^A e^{f(\mathbf{B})} [\nabla_{\mathbf{B}} f(\mathbf{B}) \otimes \nabla_{\mathbf{B}} f(\mathbf{B}) + \nabla_{\mathbf{B}} \otimes (\nabla_{\mathbf{B}} f(\mathbf{B}))] \end{aligned}$$

and therefore, using Eq. (38),

$$\mathbb{P} = n\mathbf{u} \otimes \mathbf{u} + n\nabla_{\mathbf{B}} \otimes (\nabla_{\mathbf{B}} f(\mathbf{B})). \quad (39)$$

This decomposition of  $\mathbb{P}$  shows that  $\nabla_{\mathbf{B}} \otimes (\nabla_{\mathbf{B}} f(\mathbf{B}))$  plays the role of pressure tensor in the Euler equations (34). Unfortunately, the ‘‘mass’’ tensor  $\mathbb{Q}$  has not a similarly simple expression in terms of  $f(\mathbf{B})$ .

As already remarked, the form (38) of the constraint equations allows to reduce the problem of the solvability of  $(A, \mathbf{B})$  as a function of  $(n, \mathbf{u})$  to the solvability of  $\mathbf{B}$  as a function of  $\mathbf{u}$  from the equation

$$\nabla_{\mathbf{B}} f(\mathbf{B}) = \mathbf{u},$$

which is proven in the following theorem.

**Theorem 1** *The mapping  $\mathbf{B} \in \mathbb{R}^3 \mapsto \nabla_{\mathbf{B}} f(\mathbf{B}) \in \mathbb{R}^3$  is globally invertible.*

*Proof.* We first prove local invertibility. Let  $\mathbf{u}(\mathbf{B}) := \nabla_{\mathbf{B}} f(\mathbf{B})$ . Using (39), and recalling that  $n > 0$  is given, we have that

$$\begin{aligned} \frac{\partial u_i}{\partial B_j} &= \frac{\partial^2 f}{\partial B_i \partial B_j} = \frac{\mathbb{P}_{ij}}{n} - u_i u_j \\ &= \frac{1}{n} \int_{\mathbb{R}^3} (v_i(\mathbf{p}) - u_i)(v_j(\mathbf{p}) - u_j) \phi(A, \mathbf{B}, \mathbf{p}) d\mathbf{p}, \end{aligned}$$

showing that the Jacobian matrix of the transformation is the covariance matrix of  $\mathbf{v}(\mathbf{p})$ , relative to the probability density  $\phi(A, \mathbf{B}, \mathbf{p})/n$ , which is

semi-definite positive. The positive definiteness is readily proven by direct inspection, since

$$\sum_{i,j=1}^3 \frac{\partial u_i}{\partial B_j} \xi_i \xi_j = \frac{1}{n} \int_{\mathbb{R}^3} [\boldsymbol{\xi} \cdot (\mathbf{v}(\mathbf{p}) - \mathbf{u})]^2 \phi(A, \mathbf{B}, \mathbf{p}) d\mathbf{p} > 0$$

for every  $\boldsymbol{\xi} \in \mathbb{R}^3$  with  $\boldsymbol{\xi} \neq 0$ , which concludes the proof of local invertibility.

In order to prove the global result, we resort to the classical result of Hadamard, that a local diffeomorphism is global if and only if it is proper (the inverse image of a compact is compact). In the present case this reduces to prove that, for every sequence  $\mathbf{B}_k \in \mathbb{R}^3$  such that  $|\mathbf{B}_k| \rightarrow \infty$ , also the image sequence  $\mathbf{u}_k = \mathbf{u}(\mathbf{B}_k) \in \mathbb{R}^3$  is such that  $|\mathbf{u}_k| \rightarrow \infty$ . Since  $|\mathbf{B}_k| \rightarrow \infty$ , we are interested in the asymptotic behavior of the distribution  $\phi(A, \mathbf{B}, \mathbf{p})$  for large  $|\mathbf{B}|$ . Without loss of generality, we put here  $m = 1$  and  $\beta = 1$ . The critical points of  $\phi(A, \mathbf{B}, \mathbf{p})$  (as a function of  $\mathbf{p}$ ) are determined by the condition

$$\nabla_{\mathbf{p}} (E(\mathbf{p}) - \mathbf{B} \cdot \mathbf{v}(\mathbf{p})) = 0.$$

Recalling (5) and (10), this leads to the condition

$$\mathbf{p} \pm \nabla_{\mathbf{p}} |\vec{h}(\mathbf{p})| - \mathbf{B} \mp \boldsymbol{\alpha} \cdot \mathbf{B} \nabla_{\mathbf{p}} \nu_2(\mathbf{p}) = 0,$$

that is

$$\mathbf{p} \pm \frac{(\boldsymbol{\alpha} \cdot \mathbf{p}) \boldsymbol{\alpha}}{[(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2]^{1/2}} - \mathbf{B} \mp \frac{(\boldsymbol{\alpha} \cdot \mathbf{B}) \boldsymbol{\alpha} \gamma^2}{[(\boldsymbol{\alpha} \cdot \mathbf{p})^2 + \gamma^2]^{3/2}} = 0.$$

Making the change of variable

$$\mathbf{q} = \frac{\mathbf{p}}{|\mathbf{B}|},$$

we obtain the equation

$$\mathbf{q} \pm \frac{(\boldsymbol{\alpha} \cdot \mathbf{q}) \boldsymbol{\alpha}}{|\mathbf{B}| [(\boldsymbol{\alpha} \cdot \mathbf{q})^2 + |\mathbf{B}|^{-2} \gamma^2]^{1/2}} - \frac{\mathbf{B}}{|\mathbf{B}|} \mp \frac{(\boldsymbol{\alpha} \cdot \mathbf{B}) \boldsymbol{\alpha} \gamma^2}{|\mathbf{B}|^4 [(\boldsymbol{\alpha} \cdot \mathbf{q})^2 + |\mathbf{B}|^{-2} \gamma^2]^{3/2}} = 0,$$

which is asymptotically equivalent for  $|\mathbf{B}| \rightarrow \infty$  to

$$\mathbf{q} - \frac{\mathbf{B}}{|\mathbf{B}|} = 0,$$

i.e. to

$$\mathbf{p} = \mathbf{B}.$$

Thus, we have shown that, for large  $|\mathbf{B}|$ , the distribution  $\phi(A, \mathbf{B}, \mathbf{p})$  has a single critical point (which is clearly a maximum) at  $\mathbf{p} = \mathbf{B}$ . Moreover, it decays like  $e^{-|\mathbf{p}|^2/2}$  away from the maximum. This gaussian-like behavior ensures that

$$\frac{1}{n} \int_{\mathbb{R}^3} \mathbf{p} \phi(A, \mathbf{B}, \mathbf{p}) d\mathbf{p} \sim \mathbf{B}, \quad \text{as } |\mathbf{B}| \rightarrow \infty.$$

Finally, since  $\mathbf{v}(\mathbf{p}) = \mathbf{p} \pm \nu_2(\mathbf{p})\boldsymbol{\alpha}$ , and  $\nu_2(\mathbf{p})\boldsymbol{\alpha}$  is a bounded quantity, we also obtain

$$\mathbf{u} = \frac{1}{n} \int_{\mathbb{R}^3} \mathbf{v}(\mathbf{p}) \phi(A, \mathbf{B}, \mathbf{p}) d\mathbf{p} \sim \frac{1}{n} \int_{\mathbb{R}^3} \mathbf{p} \phi(A, \mathbf{B}, \mathbf{p}) d\mathbf{p} \sim \mathbf{B},$$

which shows that  $|\mathbf{u}_k| \rightarrow \infty$  if  $|\mathbf{B}_k| \rightarrow \infty$ , concluding the proof.  $\square$

## 6 Band coupling

As already remarked, the disappearance of the quantum interference terms in the semiclassical limit makes our hydrodynamic model decoupled with respect to the two bands. Coupling mechanisms can be introduced in two ways. First of all, we may assume that the electric potential is composed of two parts:

$$V = V_{\text{ext}} + V_{\text{int}},$$

where  $V_{\text{ext}}$  is the “external” part (taking account, e.g., of external bias, gate potentials, and heterostructure potentials), while  $V_{\text{int}}$  is the “internal” (or self-consistent) part, taking account of Coulomb repulsion between electrons. In the simple mean-field model, this is given by the Poisson equation

$$\varepsilon_s \Delta V_{\text{int}} = -q(n_+ + n_-), \quad (40)$$

where  $q$  is the elementary charge and  $\varepsilon_s$  is the permittivity of the semiconductor. The right-hand side depends on the total density  $n_+ + n_-$ , this coupling the upper-band and lower-band populations.

The other source of coupling derives from collisional mechanisms. In order to introduce them, we have to go back to the kinetic level and add to the WBGK equation (18) a suitable matrix-valued “interband” collisional

operator  $C(w)$  [12]. This is assumed to act on a much slower time scale with respect to  $\tau_c$  (otherwise it would affect the hydrodynamic limit and destroy the structure of our MEP-based model). Thus, we rewrite Eq. (18) with the (generic) additional terms:

$$\begin{cases} \frac{\partial w_0}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} w_0 + \mathbf{F} \cdot \nabla_{\mathbf{p}} w_0 + \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_2 = \frac{g_0 - w_0}{\tau_c} + C_0(w), \\ \frac{\partial \vec{w}}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{x}} \vec{w} + \mathbf{F} \cdot \nabla_{\mathbf{p}} \vec{w} + \boldsymbol{\alpha} \cdot \nabla_{\mathbf{x}} w_0 \vec{e}_2 - \frac{2}{\hbar} \vec{h}(\mathbf{p}) \times \vec{w} = \frac{\vec{g} - \vec{w}}{\tau_c} + \vec{C}(w). \end{cases} \quad (41)$$

Following the same arguments that led to Eq. (29), we arrive at

$$\frac{\partial g_{\pm}}{\partial t} + v_{\pm} \cdot \nabla_{\mathbf{x}} g_{\pm} + \mathbf{F} \cdot \nabla_{\mathbf{p}} g_{\pm} = C_{\pm}(g_+, g_-) \quad (42)$$

(where we adopted a notation that stresses the fact that  $g$  only depends on  $g_+$  and  $g_-$ ). Taking the zeroth-order and first-order moments of this equation we get a modified version of the hydrodynamic system (34):

$$\begin{cases} \frac{\partial n_{\pm}}{\partial t} + \nabla_{\mathbf{x}} (n_{\pm} \mathbf{u}_{\pm}) = N_{\pm}(n_+, n_-, \mathbf{u}_+, \mathbf{u}_-), \\ \frac{\partial (n_{\pm} \mathbf{u}_{\pm})}{\partial t} + \nabla_{\mathbf{x}} \cdot \mathbb{P}_{\pm} - \mathbf{F} \cdot \mathbb{Q}_{\pm} = \mathbf{U}_{\pm}(n_+, n_-, \mathbf{u}_+, \mathbf{u}_-), \end{cases} \quad (43)$$

where, of course,

$$\begin{aligned} N_{\pm} &= \int_{\mathbb{R}^3} C_{\pm}(g_+, g_-) d\mathbf{p} \\ \mathbf{U}_{\pm} &= \int_{\mathbb{R}^3} \mathbf{v}_{\pm}(\mathbf{p}) C_{\pm}(g_+, g_-) d\mathbf{p}, \end{aligned} \quad (44)$$

and the dependence on  $(n_+, n_-, \mathbf{u}_+, \mathbf{u}_-)$  follows from the MEP closure.

Let us now list some possible choices of  $C(w)$  in a simple BGK (relaxation time) form, corresponding to different interband scattering mechanisms.

**1. Band-flip** The electron undergoes a collision which exchange its band label from  $+$  to  $-$ , or from  $-$  to  $+$ . Then we put

$$C^{bf}(w) = -\frac{w - w_0 \sigma_0}{\tau_{bf}} = -\frac{\vec{w} \cdot \vec{\sigma}}{\tau_{bf}} \quad (45)$$

(where  $\tau_{bf}$  denotes the characteristic time of band-flip scattering, which we assume constant for simplicity), so that

$$C_{\pm}^{bf}(w) = \mp \frac{w_+ - w_-}{\tau_{bf}}$$

(from which the band-flip is evident). According to definition (44), therefore, we have

$$N_{\pm}^{bf} = \mp \frac{n_+ - n_-}{\tau_{bf}}, \quad \mathbf{U}_{\pm}^{bf} = \mp \frac{n_+ \mathbf{u}_+ - n_- \mathbf{u}_-}{\tau_{bf}}. \quad (46)$$

Note that the band flip mechanism conserves the total density and the momentum and relaxes the polarization of density and momentum, (i.e.  $n_+ - n_-$  and  $\mathbf{u}_+ - \mathbf{u}_-$ ).

**2. Band relaxation** An electron in the upper band undergoes a inelastic collision which scatters it to the lower band [7]. This mechanism is described by

$$C^{br}(w) = - \frac{w_0 \vec{v} - \vec{w}}{\tau_{br}} \cdot \vec{\sigma}, \quad (47)$$

so that

$$C_{\pm}^{bf}(w) = \mp \frac{w_+}{\tau_{br}},$$

(where  $\tau_{br}$  denotes the characteristic time of band relaxation scattering, which we assume constant). From definition (44) we obtain

$$N_{\pm}^{br} = \mp \frac{n_+}{\tau_{br}}, \quad \mathbf{U}_{\pm}^{br} = \mp \frac{n_+ \mathbf{u}_+}{\tau_{br}}. \quad (48)$$

Note that this mechanism conserves the total density an momentum and depletes the upper band in favor of the lower.

**3. Isotropic interband scattering** An electron undergoes a scattering event that changes its band label and re-distributes its momentum according to a isotropic, thermal distribution. This mechanism is described by

$$C^{is}(w) = - \frac{w - g^*}{\tau_{is}}, \quad (49)$$



where  $\tau_{is}$  denotes the characteristic time of interband scattering, which we assume constant, and where  $g^*$  is the isotropic version, with inverted densities, of the MEP local equilibrium  $g$ , i.e.

$$g_{\pm}^*(\mathbf{x}, \mathbf{p}, t) = \frac{n_{\mp}}{z_{\pm}} e^{-\beta E_{\pm}(\mathbf{p})}, \quad \vec{g}_{\perp}^* = 0, \quad (50)$$

where

$$z_{\pm} = \int_{\mathbb{R}^3} e^{-\beta E_{\pm}(\mathbf{p})} d\mathbf{p}, \quad (51)$$

so that

$$\int_{\mathbb{R}^3} g_{\pm}^*(\mathbf{x}, \mathbf{p}, t) d\mathbf{p} = n_{\mp}(\mathbf{x}, t), \quad \int_{\mathbb{R}^3} \mathbf{v}_{\pm}(\mathbf{p}) g_{\pm}^*(\mathbf{x}, \mathbf{p}, t) d\mathbf{p} = 0$$

(note the inverted band-labels of the density). Then:

$$C_{\pm}^{is}(w) = -\frac{w_{\pm} - g_{\pm}^*}{\tau_{is}}$$

and

$$N_{\pm}^{is} = \mp \frac{n_{+} - n_{-}}{\tau_{is}}, \quad \mathbf{U}_{\pm}^{is} = -\frac{n_{\pm} \mathbf{u}_{\pm}}{\tau_{is}}, \quad (52)$$

Note, therefore, that this scattering mechanism relaxes the current in both bands and the density polarization .

## 7 Conclusions

We can finally summarize the hydrodynamic model emerged from our discussion. It consists of the Euler-Poisson-like system

$$\begin{cases} \frac{\partial n_{\pm}}{\partial t} + \nabla_{\mathbf{x}} \cdot (n_{\pm} \mathbf{u}_{\pm}) = N_{\pm}, \\ \frac{\partial (n_{\pm} \mathbf{u}_{\pm})}{\partial t} + \nabla_{\mathbf{x}} \cdot (n \mathbf{u}_{\pm} \otimes \mathbf{u}_{\pm} + n \mathbb{T}_{\pm}) + \nabla_{\mathbf{x}} \cdot (V_{\text{ext}} + V_{\text{int}}) \cdot \mathbb{Q}_{\pm} = \mathbf{U}_{\pm}, \\ \varepsilon_s \Delta V_{\text{int}} = -q(n_{+} + n_{-}), \end{cases} \quad (53)$$

where:

$$N_{\pm} = N_{\pm}(n_{+}, n_{-}, \mathbf{u}_{+}, \mathbf{u}_{-}), \quad \mathbf{U}_{\pm} = \mathbf{U}_{\pm}(n_{+}, n_{-}, \mathbf{u}_{+}, \mathbf{u}_{-})$$

are the coupling terms discussed above,

$$\mathbb{T}_{\pm} = \nabla_{\mathbf{B}_{\pm}} \otimes \nabla_{\mathbf{B}_{\pm}} \log \int_{\mathbb{R}^3} e^{-\beta E_{\pm}(\mathbf{p}) + \mathbf{B}_{\pm} \cdot \mathbf{v}_{\pm}(\mathbf{p})} d\mathbf{p}$$

is the pressure tensor, described in Sec. 5,

$$\mathbb{Q}_{\pm} = \int_{\mathbb{R}^3} \mathbb{M}_{\pm}^{-1}(\mathbf{p}) e^{-\beta E_{\pm}(\mathbf{p}) + \mathbf{B}_{\pm} \cdot \mathbf{v}_{\pm}(\mathbf{p}) + A_{\pm}} d\mathbf{p},$$

is the effective-mass tensor, also described in Sec. 5, and the Lagrange multipliers  $(A_{\pm}, \mathbf{B}_{\pm})$  can be uniquely solved as functions of the moments  $(n_{\pm}, \mathbf{u}_{\pm})$  from the constraint equations

$$\begin{cases} \int_{\mathbb{R}^3} e^{-\beta E_{\pm}(\mathbf{p}) + \mathbf{B}_{\pm} \cdot \mathbf{v}_{\pm}(\mathbf{p}) + A_{\pm}} d\mathbf{p} = n_{\pm}, \\ \int_{\mathbb{R}^3} \mathbf{v}_{\pm}(\mathbf{p}) e^{-\beta E_{\pm}(\mathbf{p}) + \mathbf{B}_{\pm} \cdot \mathbf{v}_{\pm}(\mathbf{p}) + A_{\pm}} d\mathbf{p} = \mathbf{u}_{\pm}, \end{cases}$$

as proven in Theorem 1.

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